

Searching for STEVE-like electron temperature spikes with Swarm

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**Special recognition to NJIT Physics 433 (Electromagnetism 2) class.

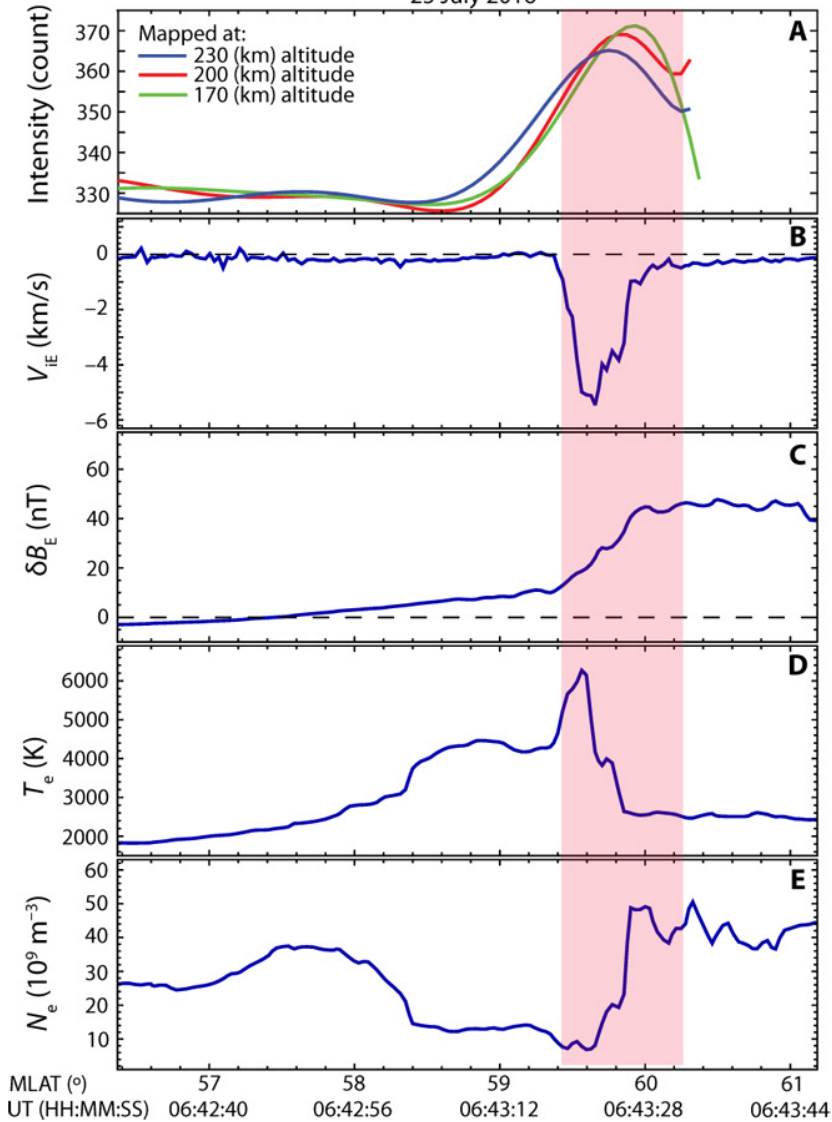
DASP Workshop 2024

February 22, 2024

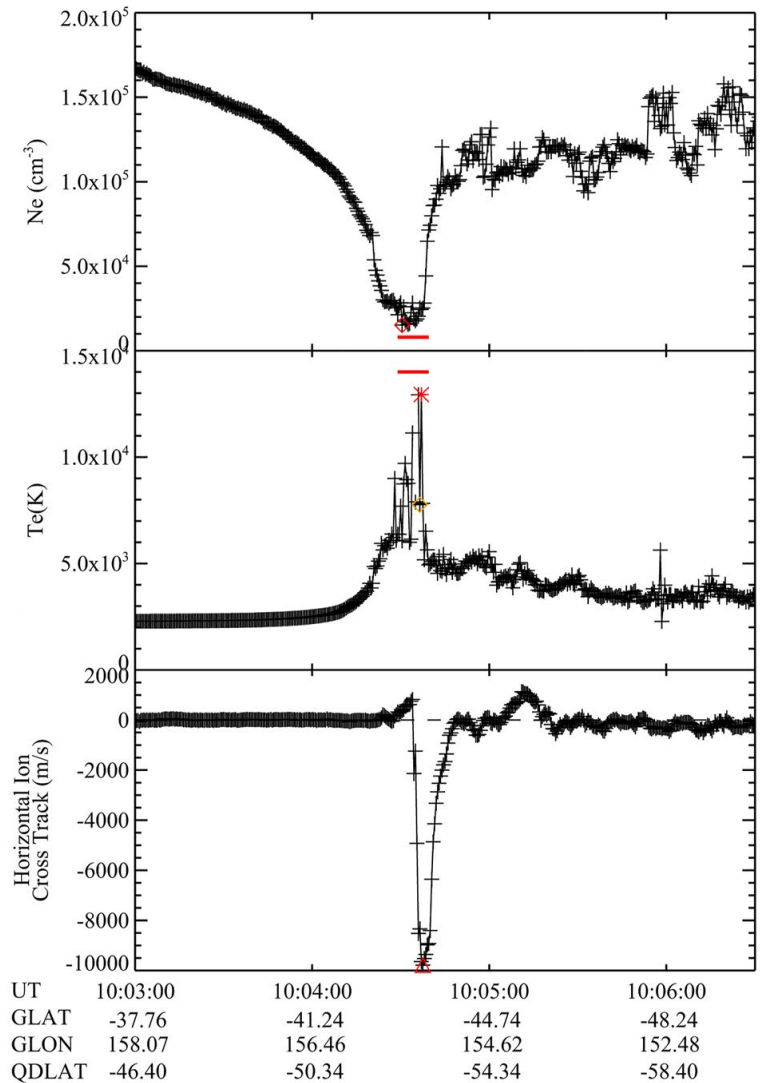
Edmonton, Alberta

This work is supported by: NSF GEM-2225972.

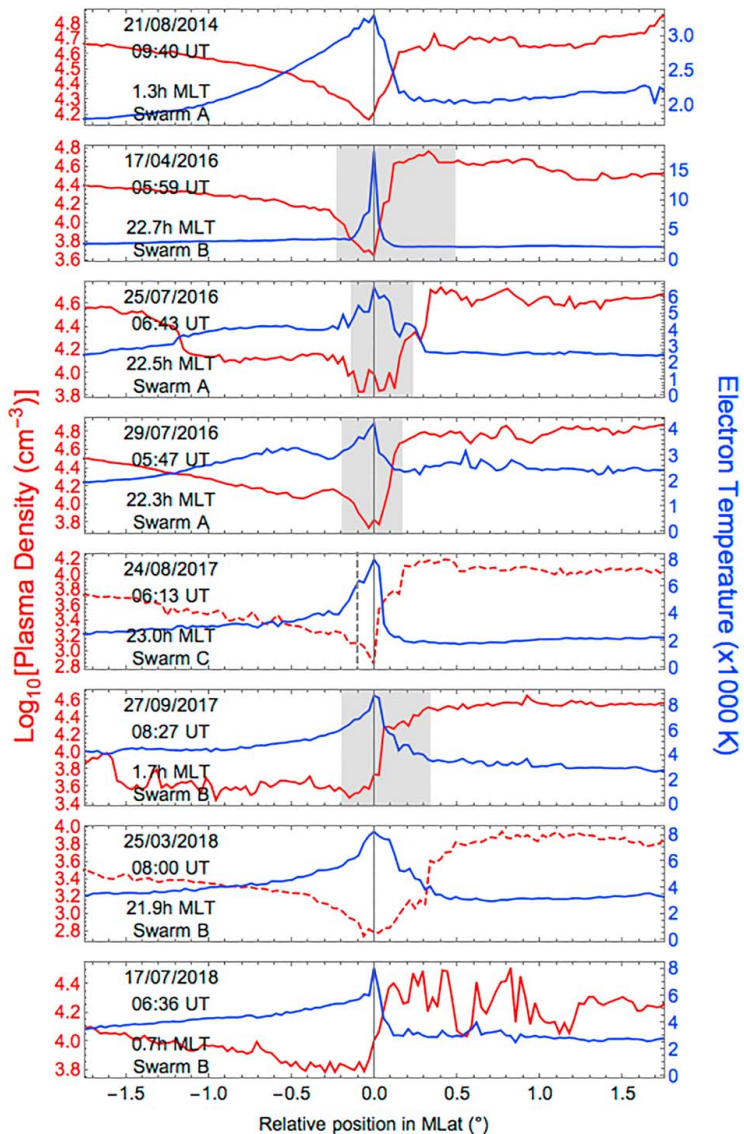
25 July 2016



(MacDonald et al., 2018)



(Martinis et al., 2022)



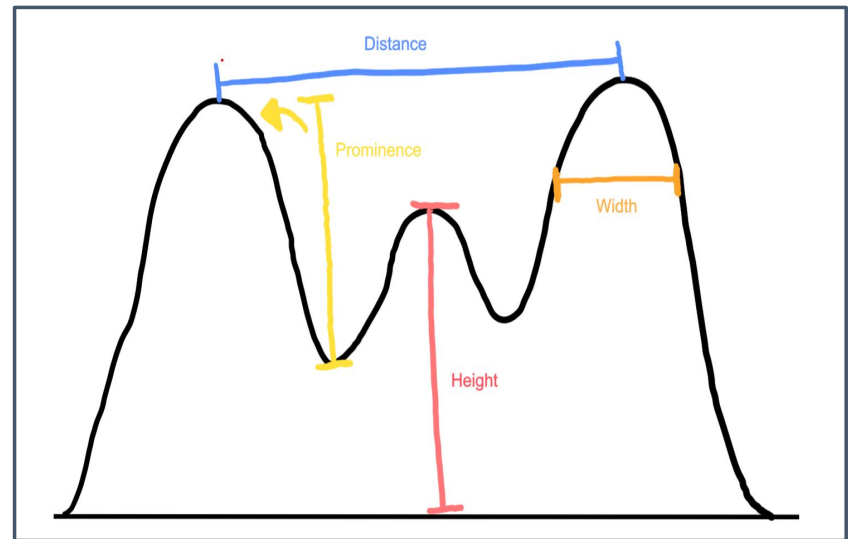
(Archer et al., 2019)

Motivation

- One of the traits of STEVE/SAID encountered by spacecraft in low Earth orbit is their prominent electron temperatures.
- Distinct optical characteristics and plasma flows are also observed, but not our focus in this presentation.
- **How common are these electron temperature “spikes”?**

Methodology

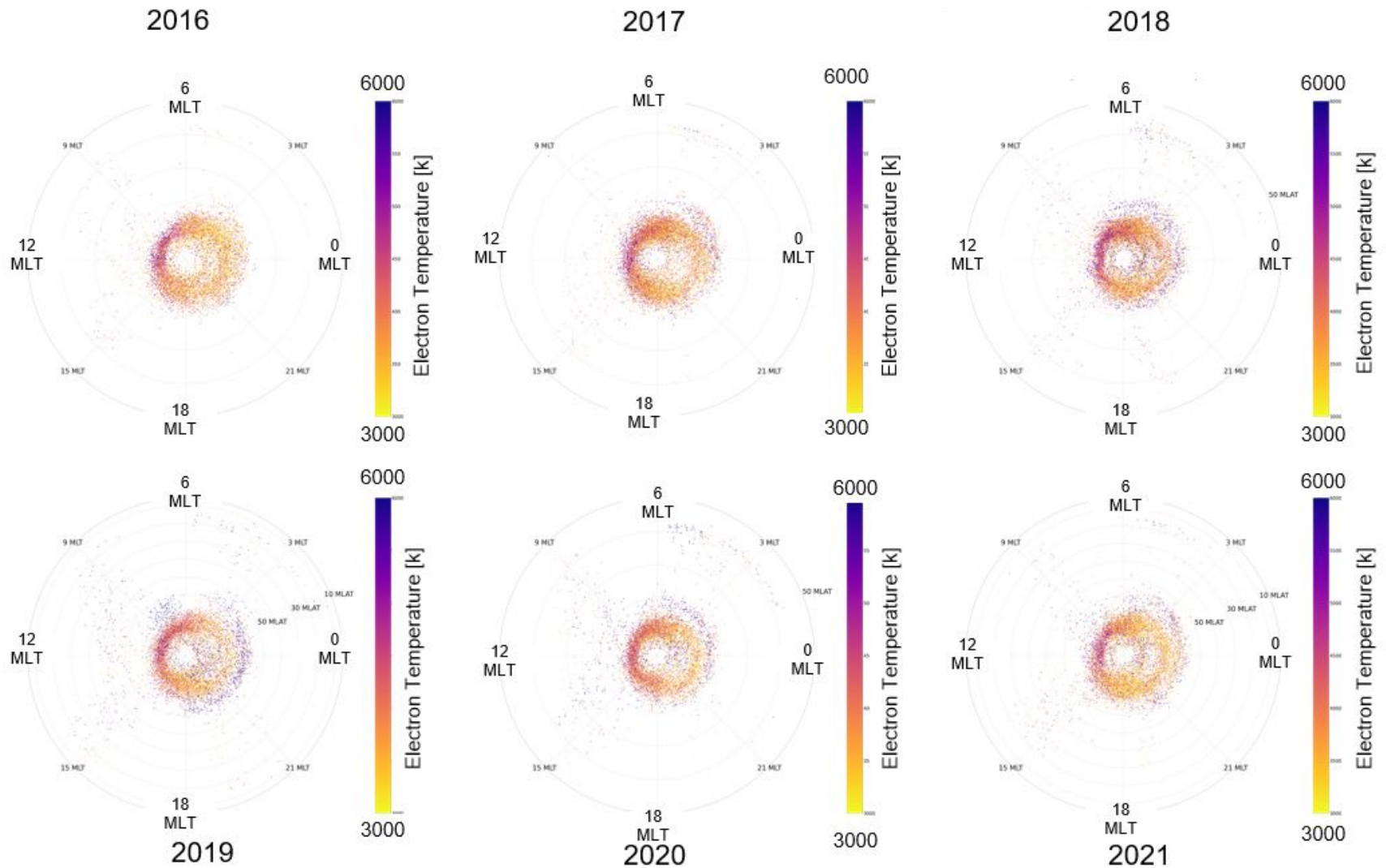
- Develop an algorithm that sifts through the Swarm electron temperature data and identifies electron temperature “spikes”.
 - Using the Swarm EFI LP data set “SW_OPER_EFIA_LP_1B”.
- Use the Archer et al. (2019) SAIDs as the archetype the spikes.
- Run the algorithm on all Swarm A, B, and C electron temperature data.



Spike conditions

- Height: 3000 - 20,000 K.
- Distance: 100 samples (~350 km along track).
- Width: 4 - 40 samples (~15 - 150 km along track).
- Prominence: 1250 - 5000 K.

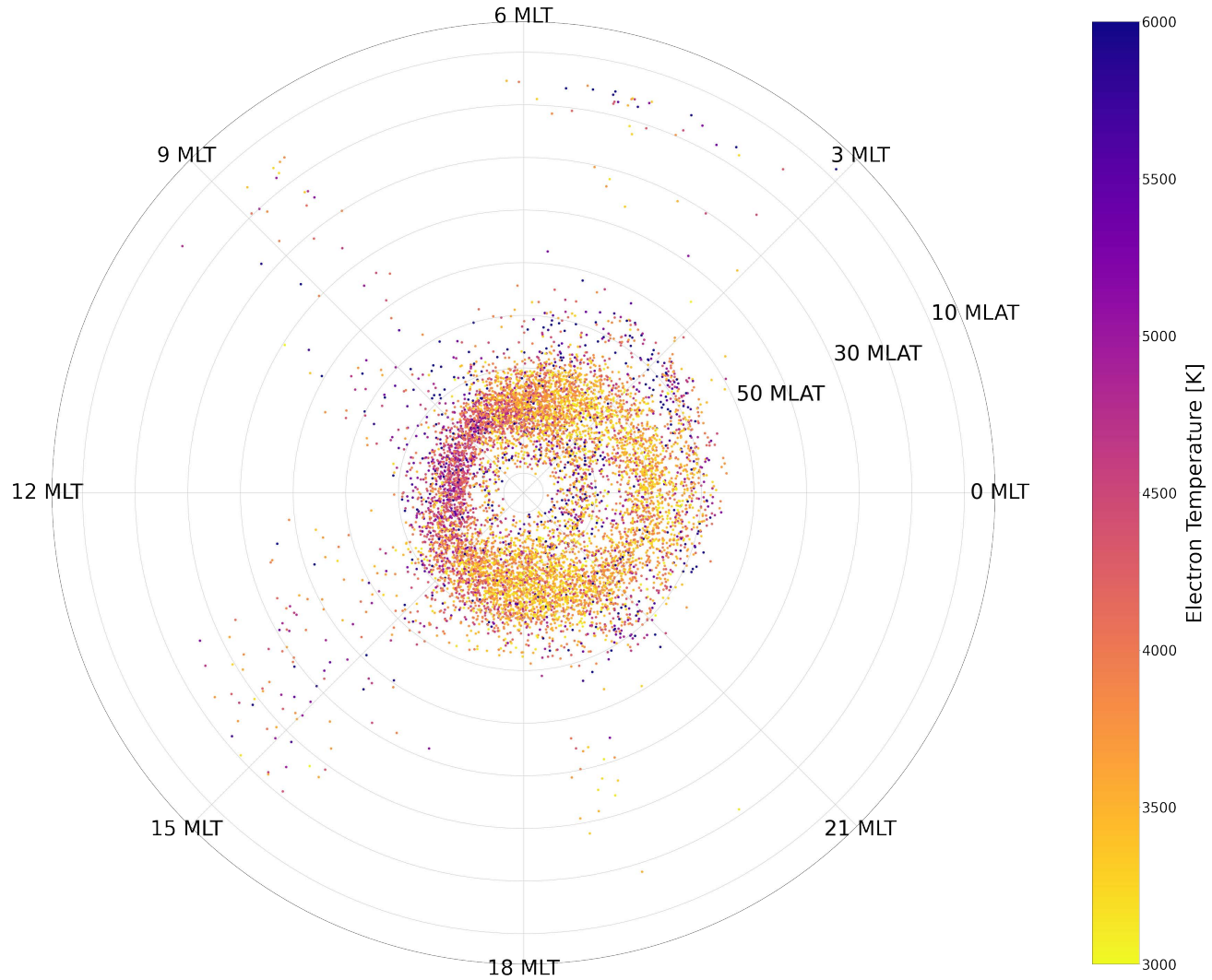
Results



Swarm A

Results

Electron spikes from North Pole to Equator



Swarm A 2021 - All spikes

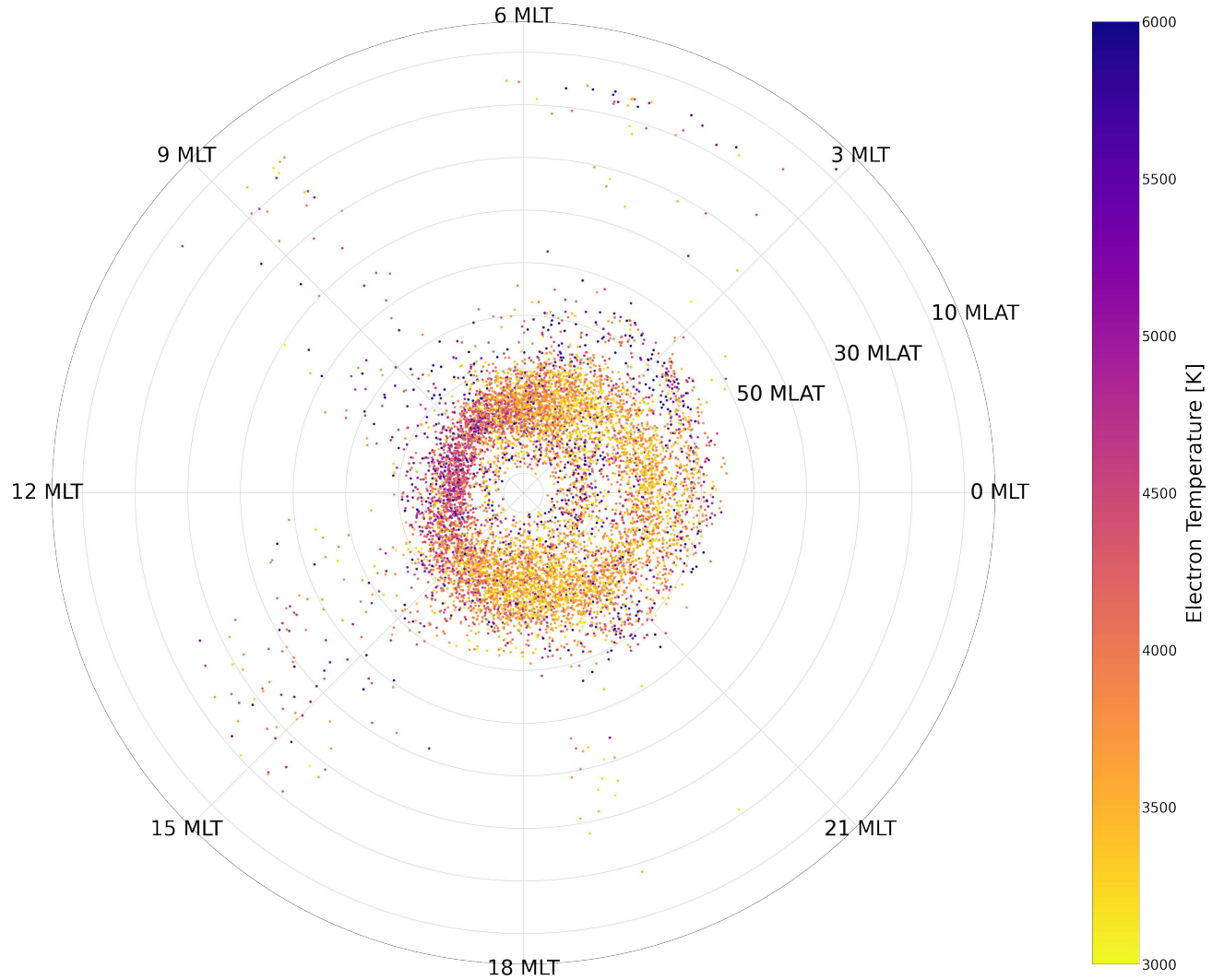
Methodology (contd..)

- To further constrain our spike events, we performed a correlation analysis to identify events with a corresponding drop in plasma density.



Results

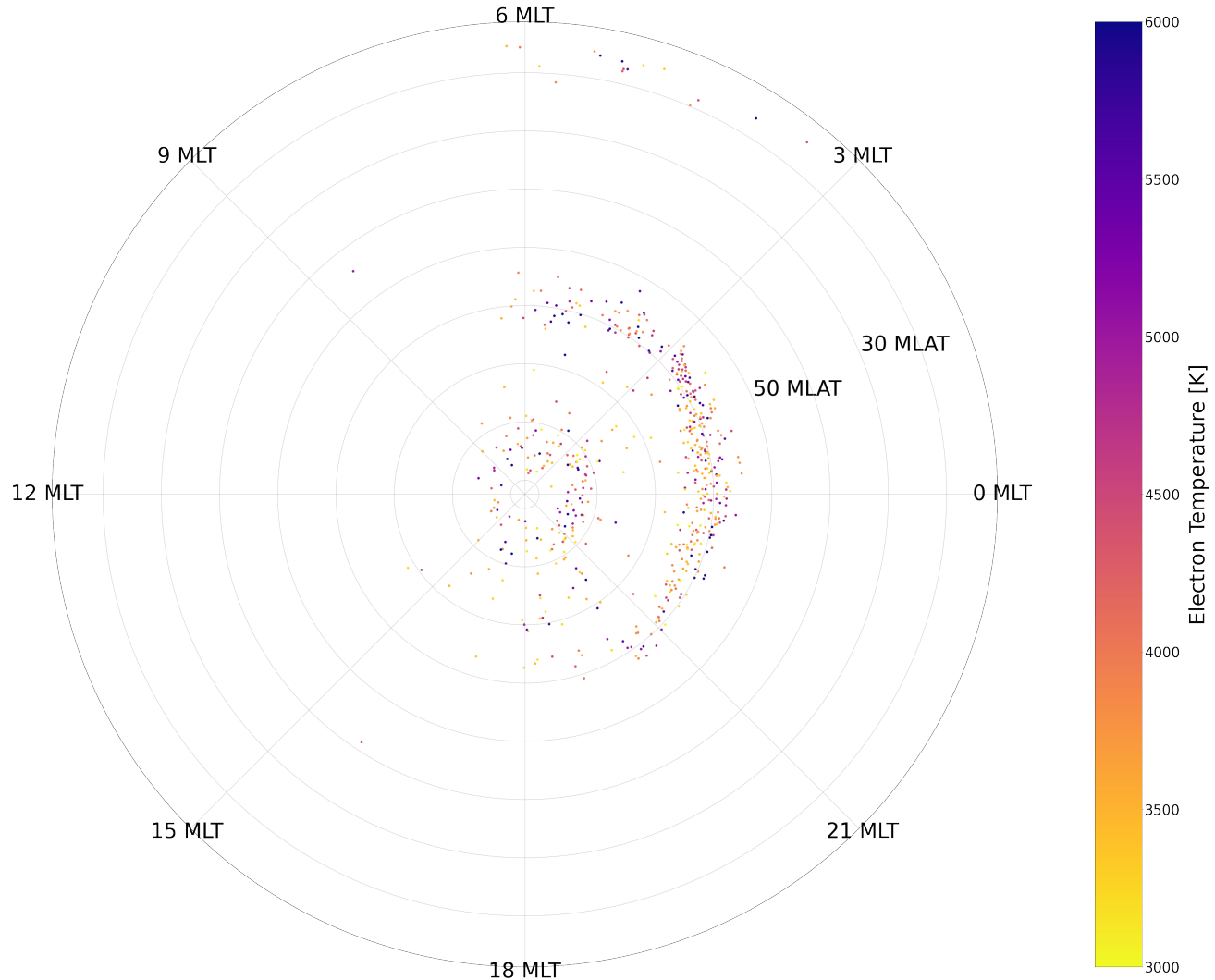
Electron spikes from North Pole to Equator



Swarm A 2021 - All spikes

Results

Electron spikes from North Pole to Equator



Swarm A 2021 - Filtered by Plasma Density Correlation

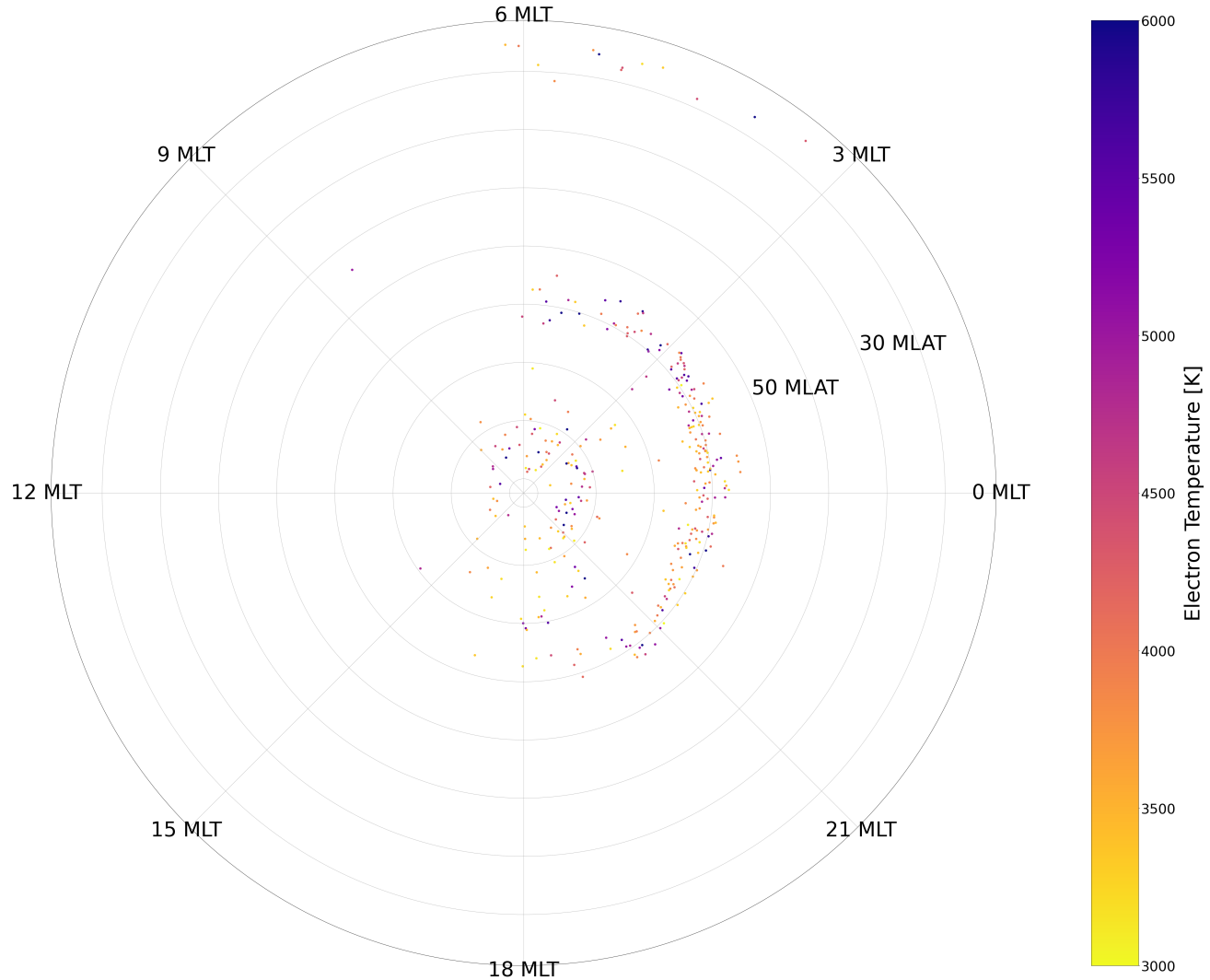
Methodology (contd..)

- After closer inspection, we realized that some of our identified events were false-positives — plasma trough crossings were being identified, so we modified the algorithm to pick out “notches” in the plasma density.



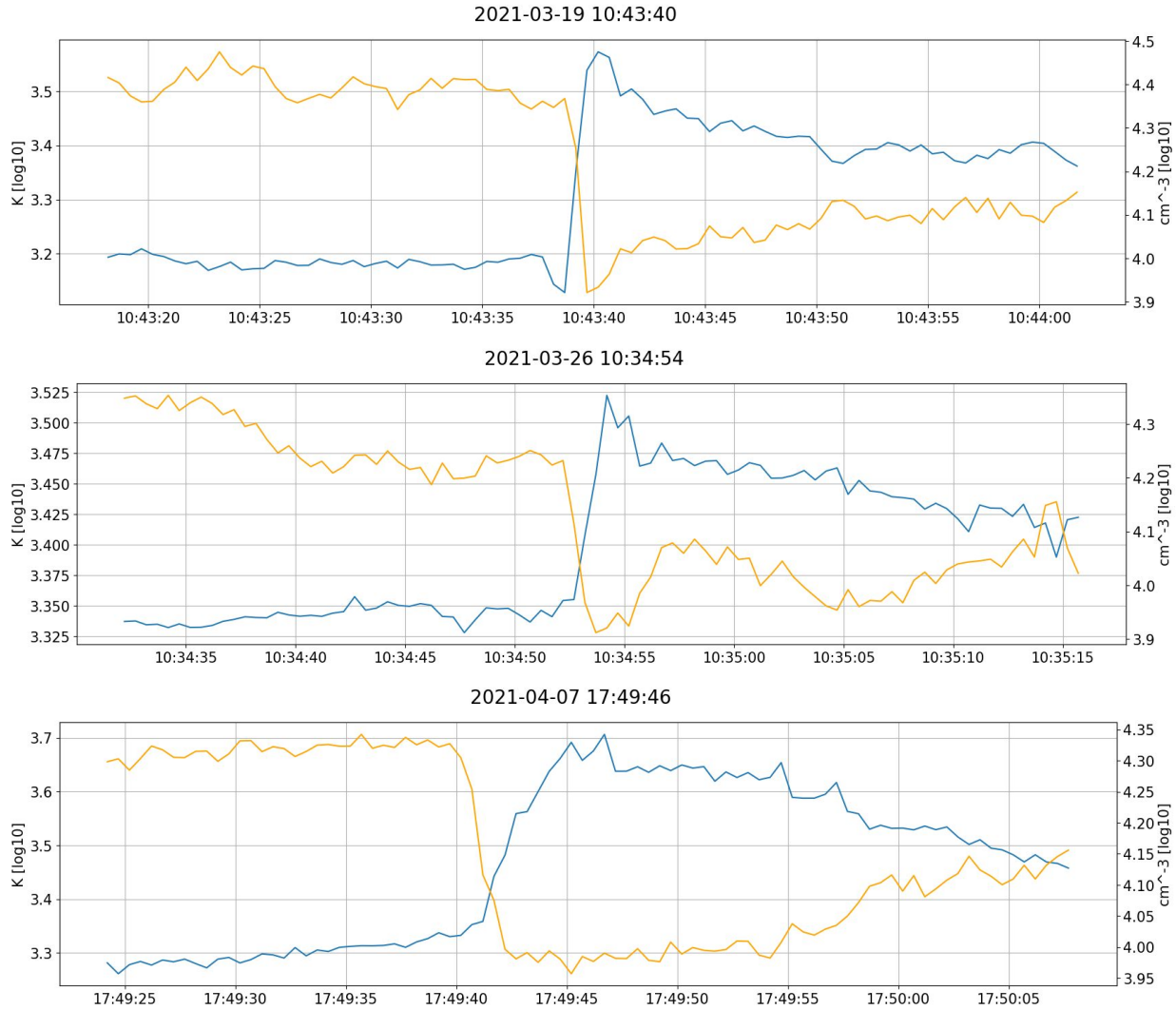
Results

Electron spikes from North Pole to Equator



Swarm A 2021 - Filtered by Plasma Density Correlation with “Drops”

Results (contd..)



Summary and Future Work

- Our motivation in this work is to determine the prevalence of STEVE-like electron temperature spikes using the Swarm data set.
 - An algorithm to perform this task was developed and implemented.
- Results show that each spacecraft observed thousands of spikes per year.
 - The spikes are organized into geomagnetic regions consistent with the auroral zone, cusp, and sub-auroral region.
- Additional analysis reveals that a subset of the spikes are associated with a plasma “trough” or “notch” decrease, and that these spikes appear to be contained to the polar-cap and sub-auroral regions.
 - It is difficult to distinguish some “trough” or “notch” plasma features from plasma trough transits.
- In future work Swarm cross-track plasma flows will be considered.

References

Archer, W. E., Gallardo-Lacourt, B., Perry, G. W., St.-Maurice, J.-P., Buchert, S. C., & Donovan, E. F. (2019). Steve: The optical signature of intense subauroral ion drifts. *Geophysical Research Letters*, 46, 6279–6286. <https://doi.org/10.1029/2019GL082687>

MacDonald, E. A., Donovan, E., Nishimura, Y., Case, N. A., Gillies, D. M., Gallardo-Lacourt, B., ... & Schofield, I. (2018). New science in plain sight: Citizen scientists lead to the discovery of optical structure in the upper atmosphere. *Science advances*, 4(3), eaaq0030

Martinis, C., Griffin, I., Gallardo-Lacourt, B., Wroten, J., Nishimura, Y., Baumgardner, J., & Knudsen, D. J. (2022). Rainbow of the night: First direct observation of a SAR arc evolving into STEVE. *Geophysical Research Letters*, 49, e2022GL098511. <https://doi.org/10.1029/2022GL098511>

Extra Slides

STEVE in ISR measurements?

2008-03-26 07:27:30 UT

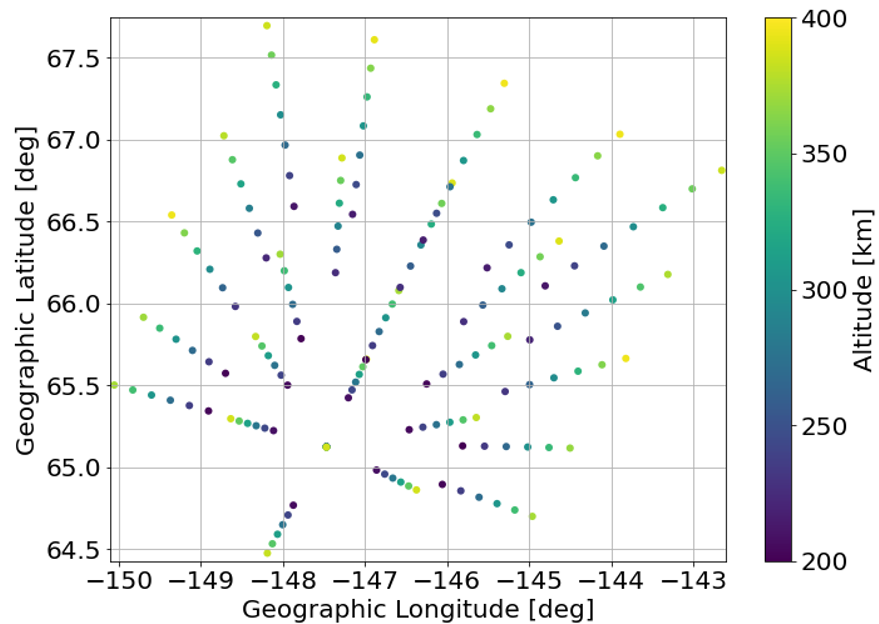
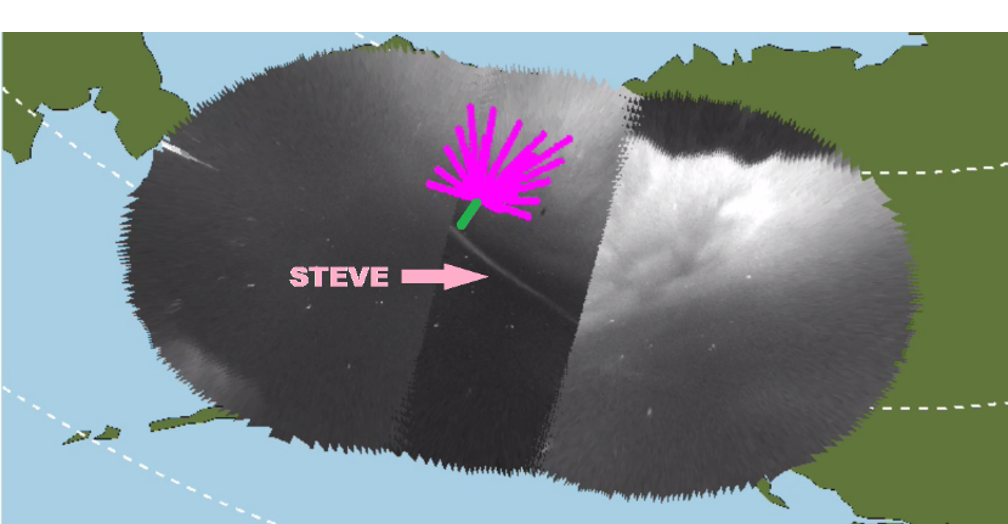


Figure courtesy of L. V. Goodwin

STEVE in ISR measurements?

**Beam 13, Azimuth = -154.3° , Elevation = 77.5°
(Antiparallel to the magnetic field)**

